



Ricci-Bourguignon Soliton on Perfect Fluid Spacetime

Mahima A., Madan M Desai, Somashekhara G., Shivaprasanna G. S., Savithri S. and Deepak A. S.

ABSTRACT: This paper examines perfect fluid spacetimes in the context of Ricci-Bourguignon solitons and η -Ricci-Bourguignon solitons, emphasizing the role of torse-forming vector fields. We derive the conditions for a Ricci-Bourguignon soliton to be shrinking, steady, or expanding, by analyzing the associated scalar parameters. Additionally, we investigate perfect fluid spacetimes satisfying the condition $(\xi, \cdot)C \cdot Z = 0$, $(\xi, \cdot)H \cdot Z = 0$.

Keywords: Ricci-Bourguignon soliton, η -Ricci-Bourguignon soliton, quasi conformal curvature tensor, conharmonic curvature tensor.

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1. Introduction

Ricci-Bourguignon Soliton (RBS) was first discovered by Jean Pierre Bourguignon in 1981 [10] and is given by

$$\frac{\partial}{\partial t}g = -2(S - \rho r g), g(0) = g_0, \quad (1.1)$$

where g is the Riemannian metric, S is the Ricci curvature tensor, r is the scalar curvature, $\rho \in \mathbb{R}$. The self-similar solutions for this flow are called Ricci-Bourguignon solitons. An n -dimensional Riemannian manifold (M, g) is said to admit RBS if

$$(\mathcal{L}_V g)(X, Y) + 2S(X, Y) + 2(\mu + \rho r)g(X, Y) = 0, \quad (1.2)$$

where $X, Y \in \chi(M)$, \mathcal{L}_V is the Lie derivative operator in the direction of vector field V , μ and ρ are constants. An n -dimensional Riemannian manifold is said to admit η -Ricci-Bourguignon soliton (η -RBS) if

$$(\mathcal{L}_V g)(X, Y) + 2S(X, Y) + 2(\mu + \rho r)g(X, Y) + 2\omega\eta(X)\eta(Y) = 0, \quad (1.3)$$

where ω is a constant.

If $\omega = 0$, η -RBS reduces to RBS.

RBS is said to be shrinking, steady or expanding according to $\mu < 0$, $\mu = 0$ or $\mu > 0$ respectively.

The stress-energy tensor or the energy-momentum tensor plays a vital role in defining the matter content of the spacetime. It describes the density, energy flux and momentum in spacetime. A Perfect Fluid

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Spacetime (P.F.S) behaves like a perfect fluid or an ideal fluid that can be completely characterized by its rest frame mass density σ and isotropic pressure p_i .

Stress-energy tensor is given by [12]

$$T(X, Y) = p_i g(X, Y) + (\sigma + p_i) \eta(X) \eta(Y), \quad (1.4)$$

for any $X, Y \in \chi(M)$, $g(X, \xi) = \eta(X)$ is the 1-form and $g(\xi, \xi) = -1$.

Einstein's Field Equation describes how the curvature of the spacetime changes by the influence of matter and energy. It is given by [12]

$$S(X, Y) + \left(\lambda - \frac{r}{2} \right) g(X, Y) = kT(X, Y), \quad (1.5)$$

where λ is the cosmological constant, $k = 8\pi G$, G being the universal gravitational constant.

Substituting (1.4) into (1.5), we obtain

$$S(X, Y) = \left(kp_i - \lambda + \frac{r}{2} \right) g(X, Y) + k(\sigma + p_i) \eta(X) \eta(Y). \quad (1.6)$$

2. Preliminaires

The (P.F.S) (M, g) satisfies (1.6). On contraction of (1.6), we get scalar curvature as

$$r = 4\lambda + k(\sigma - 3p_i). \quad (2.1)$$

Therefore, (1.6) becomes

$$S(X, Y) = \left[\lambda + \frac{k(\sigma - p_i)}{2} \right] g(X, Y) + k(\sigma + p_i) \eta(X) \eta(Y), \quad (2.2)$$

$$QX = \left[\lambda + \frac{k(\sigma - p_i)}{2} \right] X + k(\sigma + p_i) \eta(X) \xi, \quad (2.3)$$

where Q is a Ricci operator defined as $g(QX, Y) = S(X, Y)$.

If the unit timelike vector field ξ of the (P.F.S) satisfies

$$\nabla_X \xi = X + \eta(X) \xi, \quad (2.4)$$

then ξ is called torse-forming vector field.

From (2.4), for $X, Y, Z \in \chi(M)$ we can deduce the following

$$\nabla_\xi \xi = 0, \quad (2.5)$$

$$R(X, Y) \xi = \eta(Y) X - \eta(X) Y, \quad (2.6)$$

$$\eta(R(X, Y) Z) = \eta(X) g(Y, Z) - \eta(Y) g(X, Z), \quad (2.7)$$

where R is the curvature tensor of the manifold.

3. RBS on (P.F.S) with torse-forming vector field

On replacing V by ξ and substituting (2.2) in (1.2), we get

$$\left[1 + \lambda + \frac{k(\sigma + p_i)}{2} + \mu + \rho r \right] g(X, Y) + [1 + k(\sigma + p_i)] \eta(X) \eta(Y) = 0. \quad (3.1)$$

Taking $X = Y = \xi$ in the above equation, we get

$$\mu = k \left[\frac{1}{2} (\sigma + p_i) - \rho (\sigma - p_i) \right] - \lambda (1 + 4\rho). \quad (3.2)$$

Therefore, we can state the following theorem.

Theorem 3.1 *Let (P.F.S) admitting the RBS with the torse-forming vector field ξ . Then, the soliton is said to be shrinking or steady or expanding as $k \left[\frac{1}{2}(\sigma + p_i) - \rho(\sigma - p_i) - \lambda(1 + 4\rho) \right] < 0$ or $k \left[\frac{1}{2}(\sigma + p_i) - \rho(\sigma - p_i) - \lambda(1 + 4\rho) \right] = 0$ or $k \left[\frac{1}{2}(\sigma + p_i) - \rho(\sigma - p_i) - \lambda(1 + 4\rho) \right] > 0$.*

If

$$(\mathcal{L}_V g)(X, Y) = 2\Phi g(X, Y), \quad (3.3)$$

then V is called the conformal Killing vector field.

V is called proper if Φ is a non-constant. V is called a homothetic vector field if $\Phi \neq 0$ is a constant. Consider V to be the conformal Killing vector field in (1.2). Then from (1.2) and (3.3) we get

$$S(X, Y) = -[\Phi + \mu + \rho r]g(X, Y). \quad (3.4)$$

Therefore, the spacetime is Einstein.

Conversely, if the spacetime is Einstein (i.e., $S(X, Y) = \theta g(X, Y)$) then (1.2) is reduced to

$$(\mathcal{L}_V g)(X, Y) = 2\Psi g(X, Y), \quad (3.5)$$

where $\Psi = -(\theta - \mu - \rho r)$.

As a result, we can state that

Theorem 3.2 *The potential vector field V in (P.F.S) admitting RBS is conformal Killing vector field if and only if the spacetime is Einstein.*

By virtue of (2.2) and (3.4) we obtain

$$\left[\Phi + \mu + \rho r + \lambda + \frac{k}{2}(\sigma - p_i) \right] g(X, Y) + k(\sigma + p_i)\eta(X)\eta(Y) = 0. \quad (3.6)$$

Taking $Y = \xi$, we get

$$\left[\Phi + \mu + \rho r + \lambda + \frac{k}{2}(\sigma - p_i) - k(\sigma + p_i) \right] \eta(X) = 0. \quad (3.7)$$

Since, 1-form $\eta(X) \neq 0$, we have

$$\Phi = kp_i(1 + 2\rho). \quad (3.8)$$

We obtain Φ as a constant. Hence, we can state the theorem as follows:

Theorem 3.3 *If (PFS) admits RBS with the torse-forming vector field and the potential vector field V be the conformal Killing vector field, then V is the homothetic vector field.*

4. η -RBS on (P.F.S)

Let (M^4, g) be (P.F.S) and (g, μ, ω, ξ) be the η -RBS on M^4 . Then by the equations (2.1), (2.2) and (1.3) we have

$$\frac{1}{2}(\mathcal{L}_V g)(X, Y) + \left\{ \lambda + \frac{k(\sigma - p_i)}{2} + \mu + \rho[4\lambda + k(\sigma - 3p_i)] \right\} g(X, Y) + [k(\sigma + p_i) + \omega]\eta(X)\eta(Y) = 0. \quad (4.1)$$

Let $e_i (1 \leq i \leq 4)$ be an orthonormal frame field. Therefore, $\xi = \sum_{i=1}^4 \xi^i e_i$ and from [2], $\sum_{i=1}^4 \epsilon_{ii}(\xi^i)^2 = -1$ and $\eta(e_i) = \epsilon_{ii}\xi^i$. Multiply (4.1) by ϵ_{ii} and sum over i for $X = Y = e_i$ to get

$$4\mu - \omega = (-4\lambda - k\sigma + 3kp_i)(4\rho + 1) - \text{div}(\xi). \quad (4.2)$$

Lie derivative in the direction of V is given by

$$(\mathcal{L}_V g)(X, Y) = \frac{1}{2}[g(\nabla_X V, Y) + g(X, \nabla_Y V)]. \quad (4.3)$$

Using (1.3),(4.1) and (2.2), we obtain

$$\mu - \omega = \lambda(1 + 4\rho) + k\sigma(1 - \rho) + 3kp_i \left(\frac{1}{2} + \rho \right). \quad (4.4)$$

Solve (4.2) and (4.4) simultaneously to get

$$\mu = \frac{1}{4}(-4\lambda - k\sigma + 3kp_i)(4\rho + 1) + \frac{k}{12}(5\sigma + 3p_i) - \frac{\text{div}(\xi)}{12}, \quad (4.5)$$

$$\omega = -\frac{k}{3}(5\sigma + 3p_i) - \frac{\text{div}(\xi)}{3}. \quad (4.6)$$

Hence, we conclude that

Theorem 4.1 *Let (M^4, g) be a Riemannian manifold and η be the g -dual 1-form of the gradient vector field $\xi = \text{grad}(f)$, with $g(\xi, \xi) = -1$, where f is a smooth function. If (g, μ, ω, ξ) is the η -RBS on M^4 , then Laplacian equation satisfied by f in η -RBS becomes $\Delta(f) = 12\mu - 3(-4\lambda - k\sigma + 3kp_i)(4\rho + 1) + k(5\sigma + 3p_i)$.*

Definition 1 [5] studied the Z -tensor which is given as

$$Z(X, Y) = S(X, Y) + \phi g(X, Y), \quad (4.7)$$

where ϕ is arbitrary scalar function.

5. (P.F.S) satisfying $(\xi, \cdot)C \cdot Z = 0$

A Quasi Conformal curvature tensor C on a differentiable manifold (M^4, g) with Riemannian metric g is given by

$$\begin{aligned} C(X, Y)Z &= aR(X, Y)Z + b[S(Y, Z)X - S(X, Z)Y + g(Y, Z)QX \\ &\quad - g(X, Z)QY] - \frac{r}{4} \left(\frac{a}{3} + 2b \right) [g(Y, Z)X - g(X, Z)Y], \end{aligned} \quad (5.1)$$

where $a, b \in \mathbb{R}$. [6], In (P.F.S) the condition $(\xi, \cdot)C \cdot Z = 0$ is equivalent to

$$Z(C(\xi, X)Y, \xi) + Z(Y, C(\xi, X)\xi) = 0. \quad (5.2)$$

By equations (4.7) and (5.2), we get

$$S(C(\xi, X)Y, \xi) + S(X, C(\xi, X)\xi) + \phi[\eta(C(\xi, X)Y) + g(Y, C(\xi, X)\xi)] = 0. \quad (5.3)$$

By virtue of (2.2),(2.6) and (2.7), we obtain

$$k(\sigma + p_i) \left\{ a + 2b \left[\lambda + \frac{k(\sigma - p_i)}{2} \right] - b[k(\sigma + p_i)] - \frac{r}{4} \left(\frac{a}{3} + 2b \right) \right\} [g(X, Y) + \eta(X)\eta(Y)] = 0. \quad (5.4)$$

Since $[g(X, Y) + \eta(X)\eta(Y)] \neq 0$, we have

$$p_i = \frac{\sigma(a + 6b)}{(a - 18b)} - \frac{4a(3 - \lambda)}{k(a - 18b)}. \quad (5.5)$$

The theorem can be stated as

Theorem 5.1 *Let (M^4, g) be a (P.F.S) which satisfies $(\xi, \cdot)C \cdot Z = 0$ where C is Quasi conformal curvature tensor and ξ is torse-forming vector field. Then*

$$p_i = \frac{\sigma(a + 6b)}{(a - 18b)} - \frac{4a(3 - \lambda)}{k(a - 18b)}.$$

6. (P.F.S) satisfying $(\xi, \cdot)H \cdot Z = 0$

A Conharmonic curvature tensor H on a differentiable manifold (M^4, g) with Riemannian metric g is given by

$$H(X, Y)Z = R(X, Y)Z - \frac{1}{2}[S(Y, Z)X - S(X, Z)Y + g(Y, Z)QX - g(X, Z)QY]. \quad (6.1)$$

In (P.F.S) the condition $(\xi, \cdot)C \cdot Z = 0$ is equivalent to

$$Z(H(\xi, X)Y, \xi) + Z(Y, H(\xi, X)\xi) = 0. \quad (6.2)$$

By equations (4.7) and (6.2), we get

$$S(H(\xi, X)Y, \xi) + S(X, H(\xi, X)\xi) + \phi[\eta(H(\xi, X)Y) + g(Y, H(\xi, X)\xi)] = 0. \quad (6.3)$$

By virtue of (2.2),(2.6) and (2.7), we have

$$k(\sigma + p_i) \left\{ 1 - \frac{1}{2} \left[\lambda + \frac{k(\sigma - p_i)}{2} \right] + \frac{k(\sigma + p_i)}{4} \right\} [g(X, Y) + \eta(X)\eta(Y)] = 0. \quad (6.4)$$

Since $[g(X, Y) + \eta(X)\eta(Y)] \neq 0$, we obtain

$$p_i = \frac{\lambda - 2}{k}. \quad (6.5)$$

The theorem can be stated as

Theorem 6.1 *Let (M^4, g) be (P.F.S) which satisfies $(\xi, \cdot)H \cdot Z = 0$ where H is Conharmonic curvature tensor and ξ is torse-forming vector field. Then*

$$p_i = \frac{\lambda - 2}{k}.$$

7. (P.F.S) satisfying $(\xi, \cdot)Z \cdot C = 0$

In (P.F.S), the condition $(\xi, \cdot)Z \cdot C = 0$ is equivalent to

$$\begin{aligned} & Z(X_1, C(X_2, X_3)X_4)\xi - Z(\xi, C(X_2, X_3)X_4)X_1 \\ & + Z(X_1, X_2)C(\xi, X_3)X_4 - Z(\xi, X_2)C(X_1, X_3)X_4 \\ & + Z(X_1, X_3)C(X_2, \xi)X_4 - Z(\xi, X_3)C(X_2, X_1)X_4 \\ & + Z(X_1, X_4)C(X_2, X_3)\xi - Z(\xi, X_4)C(X_2, X_3)X_1 = 0, \end{aligned} \quad (7.1)$$

where $X_1, X_2, X_3, X_4 \in \chi(M)$.

Contracting the above equation to ξ , we have

$$\begin{aligned} & -Z(X_1, C(X_2, X_3)X_4) - Z(\xi, C(X_2, X_3)X_4)\eta(X_1) \\ & + Z(X_1, X_2)\eta(C(\xi, X_3)X_4) - Z(\xi, X_2)\eta(C(X_1, X_3)X_4) \\ & + Z(X_1, X_3)\eta(C(X_2, \xi)X_4) - Z(\xi, X_3)\eta(C(X_2, X_1)X_4) \\ & + Z(X_1, X_4)\eta(C(X_2, X_3)\xi) - Z(\xi, X_4)\eta(C(X_2, X_3)X_1) = 0. \end{aligned} \quad (7.2)$$

Taking $X_3 = X_4 = \xi$ and by the virtue of (2.2) (2.3),(2.6),(2.7), (4.7),(5.1) and (7.2), we get

$$[g(X_1, X_2) + \eta(X_1)\eta(X_2)][-a - 2bL + bK + R][-L - \phi] = 0, \quad (7.3)$$

where $L = \left[\lambda + \frac{k(\sigma - p_i)}{2} \right]$, $K = k(\sigma + p_i)$ and $R = \frac{r}{4} \left(\frac{a}{3} + 2b \right)$.

Since $[g(X_1, X_2) + \eta(X_1)\eta(X_2)] \neq 0$, we get

$$p_i = \frac{4a + 4\lambda(2b - 1) - k\sigma}{8b - 3}.$$

Therefore, we can state the following theorem as

Theorem 7.1 *Let (M^4, g) be a (P.F.S) satisfying $(\xi, \cdot)Z \cdot C = 0$ with the torse-forming vector field ξ . Then $p_i = \frac{4a + 4\lambda(2b - 1) - k\sigma}{8b - 3}$.*

8. (P.F.S) satisfying $(\xi, \cdot)Z \cdot H = 0$

In (P.F.S), the condition $(\xi, \cdot)Z \cdot H = 0$ is equivalent to

$$\begin{aligned} & Z(X_1, H(X_2, X_3)X_4)\xi - Z(\xi, H(X_2, X_3)X_4)X_1 \\ & + Z(X_1, X_2)H(\xi, X_3)X_4 - Z(\xi, X_2)H(X_1, X_3)X_4 \\ & + Z(X_1, X_3)H(X_2, \xi)X_4 - Z(\xi, X_3)H(X_2, X_1)X_4 \\ & + Z(X_1, X_4)H(X_2, X_3)\xi - Z(\xi, X_4)H(X_2, X_3)X_1 = 0, \end{aligned} \quad (8.1)$$

where $X_1, X_2, X_3, X_4 \in \chi(M)$.

Contracting the above equation to ξ , we have

$$\begin{aligned} & -Z(X_1, H(X_2, X_3)X_4) - Z(\xi, H(X_2, X_3)X_4)\eta(X_1) \\ & + Z(X_1, X_2)\eta(H(\xi, X_3)X_4) - Z(\xi, X_2)\eta(H(X_1, X_3)X_4) \\ & + Z(X_1, X_3)\eta(H(X_2, \xi)X_4) - Z(\xi, X_3)\eta(H(X_2, X_1)X_4) \\ & + Z(X_1, X_4)\eta(H(X_2, X_3)\xi) - Z(\xi, X_4)\eta(H(X_2, X_3)X_1) = 0. \end{aligned} \quad (8.2)$$

Taking $X_3 = X_4 = \xi$ and by the virtue of (2.2) (2.3), (2.6), (2.7), (4.7), (6.1) and (8.2), we get

$$[g(X_1, X_2) + \eta(X_1)\eta(X_2)] \left[-1 - L + \frac{K}{2} \right] [-L - \phi] = 0, \quad (8.3)$$

where $L = \left[\lambda + \frac{k(\sigma - p_i)}{2} \right]$, $K = k(\sigma + p_i)$ and $R = \frac{r}{4} \left(\frac{a}{3} + 2b \right)$.

Since $[g(X_1, X_2) + \eta(X_1)\eta(X_2)] \neq 0$, we get

$$p_i = 4 - k\sigma.$$

Therefore, we can state the following theorem as

Theorem 8.1 *Let (M^4, g) be a (P.F.S) satisfying $(\xi, \cdot)Z \cdot H = 0$ with the torse-forming vector field ξ . Then $p_i = 4 - k\sigma$.*

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Mahima A., Somashekhara G. and Madan M Desai,
 Department of Mathematics and Statistics,
 M.S.Ramaiah University of Applied Sciences,
 Bengaluru, India.

E-mail address: mahimamahi20012003@gmail.com, somashekhara.mt.ns@msruas.ac.in, madan.mp@msruas.ac.in

ORCID: <http://orcid.org/0009-0001-9188-3405>,

<http://orcid.org/0000-0002-9069-8814>, <http://orcid.org/0009-0004-8661-0244>

and

Shivaprasanna G. S.,
 Department of Mathematics,
 Dr.Ambedkar institute of technology,
 Bengaluru, India.

E-mail address: shivaprasanna28@gmail.com

ORCID: <http://orcid.org/0000-0002-9962-5216>

and

Savithri S.,
 Department of Mathematics,
 R.V.College of Engineering,
 Bengaluru, India.

E-mail address: savithrishashidhar@rvce.edu.in

ORCID: <http://orcid.org/0000-0003-1501-4416>

and

Deepak A. S.,
 T A Pai Management Institute, Bengaluru,
 Manipal Academy of Higher Education,
 Manipal, India.

E-mail address: deepak.as@manipal.edu

ORCID: <http://orcid.org/0000-0001-5890-1512>